(27) LR + LC.T. = LBARE

finite singular Singular but unphysical but unphysical but physical but physical but physical but physical in the sense of arbitrary separation of finite terms

Tand notice that the physical results should not depend on how we shuffle the finite terms around!

and, by construction, here in Eq. (27) delivers finite correlator.

Aside: one can (not must!) choose a scheme where LR is constructed in terms of physical "couplings, e.g. M2->2 as measured at some Kinentical coinf. would allows us in ϕ^{ξ} theory to here $d_{R} = -\lambda_{R}\phi^{\xi}$ with $H_{2\rightarrow 2}(\bar{\epsilon} = \bar{\epsilon}) = -\lambda_{R}$. Or one can renormalite the mass to the physical one, ma= mphys. While this is certainly or, it's just an arbitrary choice out of infinitely many that one equally relial, (such as e.g. picking he at some unphysical Kinemetic of Mz-z). Moreover, it would just remains as orbitrers as Keeping around the "u" in this reg, since one would have E as new powereth, and centering physics does not depend on whether we decide to define the coupling as $N_2 \rightarrow 2(E=\overline{E})$ rather than to some other E=E. In other words, M = > E, no much of a difference then! And in fact it's often very useful, as we will see, not to stick to "physical" choice of the renormalited couplings or masses, especially because it leads to premeture break down of perturbetion theory due to large logarthm (see Weinbay ch. 18 for an account) e.g. $M_{2\rightarrow 2}(E=\overline{E})=\lambda_R \Rightarrow M_{2\rightarrow 2}(E)=-\lambda(1+\frac{\lambda_R}{L_{EM}})$, the record ten becoming larger than tree-level - IR for large la(E/E) even if he as this can be solved boline vie RG-equation epplied to $\mu = \overline{\epsilon}$ and recalling that $\lambda_R = \lambda(\overline{\epsilon}) = \lambda(\mu)$ and verying $\overline{\epsilon} = \mu$

- Exemple:
$$\mathcal{L}_{p} = (\partial \phi)^{2} + m^{2} \phi^{2} + \frac{\lambda}{4!} \phi^{4} \mu^{4-0} = in dim-reg.$$

(28)
$$\lambda_{c.7} = \frac{6264}{2}(4)^2 + \frac{5m^2m^2}{2} + \frac{51\mu^6}{4!} \phi^4$$

(3)
$$\int_{e^{4}} \int_{e_{7}} = \frac{1+\sqrt{2}}{2} \left(\frac{1+\sqrt{2}}{2} \right) \left(\frac{1+\sqrt{2$$

(30)
$$\Rightarrow$$
 $\mathcal{L}_{BARE} = (2 + 0)^2 + m_0^2 + \frac{1}{4!} + \frac{1}{4!} + \frac{1}{4!} + \frac{1}{4!}$ with

We can perform equivelent splittings of the same, given, Lesse

es physical results do not depend on how we shuffle the finite terms (at fixed Lacre, which "defines" the theory).

In dim-reg the freedom of the cloice of finite terms is represented by the cloice of μ (see eg. $H_{z\to z} = -\lambda \left(1 + \frac{\lambda}{2\pi i} \exp(1)\right)$ where changing $\mu - \nu \mu e^{\alpha}$ is like redefining $\lambda_R = \lambda$ by the finite shift $\lambda_R = \nu \lambda_R + \frac{2\pi i}{4\pi i} \lambda_R^2 \alpha$)

Teside: in a different scheme, e.g. where we defined λ vie H_2 - $\pi_2(\varepsilon=\overline{\varepsilon})=\lambda_R$, this fraction of changing the finite terms is essociated to changing the $\overline{\varepsilon}$ where one defines the λ_R . In the hond cutoff core one subtract lay Λ^2/μ^2 and the undetermined finite constant μ is back. It is necessarily so become c.T. are defined only up to finite terms.

(32)
$$\mathcal{L}_{R} + \mathcal{L}_{C.T.} = \mathcal{L}_{R} + \mathcal{L}_{C.T} \left(=\mathcal{L}_{RARE}\right)$$
 (Life in Wilsonian RG identif. A with \mathcal{L}_{R}) identif. A with \mathcal{L}_{R}

in dim-Rey Le= Le(\(\lambda\lambda\mu); Le.T(\(\lambda\lambda\mu)\) (+to compensate K=\(\mu\))

we calculated in pt in D=4 et 1-loop in dim-reg -

 $M_{2-02} = -\lambda \left[1 + \frac{\lambda}{32\pi^2} \left(38_{\epsilon} - 6 + h \frac{stu}{(4\pi^{-1})^3} - i\pi\right] + ...$ (33) $\sim -\lambda - \frac{6\lambda^2}{32\pi^2} \log E$

(Those constants too can be absorbed in $\mu \to \mu$) Since physics should be independent of $\mu \to M_2 - 1$ com't depend separately on λ & μ . (In fact, working to this 1-loop order it depends only on the combination $\hat{\lambda} = \lambda - 6 \frac{12}{32\pi^2} \log \mu$, $M_2 = -\hat{\lambda} - 6 \hat{\lambda}^2 \ln \xi$)

So we see that (), n) is a redundant pair!

1=1(p) such that of Obs =0 (egain like vilsanian Rg. Avilsan-op)

frexample:

 $\frac{d}{dk_{2}} = 0 = 0 + \frac{d\lambda}{dk_{1}} = \frac{6\lambda^{2}}{32\pi^{2}} = \frac{3}{16\pi^{2}}\lambda^{2} = 13\phi^{4}$

In enother sclene, e.g. $\lambda=\lambda(\overline{\epsilon})$ defined by $M_{2}=2(\overline{\epsilon})=-\lambda(1+\frac{\lambda}{32\pi i^2}\log\frac{4\pi}{\overline{\epsilon}})$, we see that $(\lambda, \bar{\epsilon})$ is not on indep. pair, menifoldy, and d cos = 0 such as $\bar{\epsilon}q - (34) - 0$ for $\bar{\epsilon}M_{2} - 2 = 0$ => $-\beta + \frac{3}{16\pi^{2}} = 0$ =K] We could have obtained the same domending that the bare parameters, i.e. those that define LBARE (that define the theory at hand) one u-independent

Example:
$$d^4$$
 of one-loop in $D=4-\epsilon$

(35)
$$\lambda_0 = \mu^{\epsilon} \left(\frac{1+d^2}{2} \right)^2 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} \right) \Rightarrow \lambda_0 = \mu^{\epsilon} \lambda \left(1 + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon}$$

(36)
$$\mu \frac{d\lambda_0}{dh_{1/2}} = \epsilon \lambda \left(1 + \frac{3\lambda}{16\pi^2} \frac{1}{\epsilon}\right) + \beta \left(1 + \frac{3\lambda}{16\pi^2} \frac{1}{\epsilon}\right) + 3 \frac{1}{16\pi^2} \frac{1}{\epsilon} = 0$$

(37)
$$\Rightarrow$$
 $\beta \left(1 + \frac{6\lambda}{16\pi^2} + \frac{1}{6\pi^2}\right) + \epsilon \lambda + \frac{3\lambda^2}{16\pi^2} = 0 \Rightarrow b \text{ lowest order } \beta = -\epsilon \lambda$

Plug backing

(38)
$$\beta - \frac{6\lambda^2}{16\pi^2} + \epsilon\lambda + \frac{3\lambda^2}{16\pi^2} = 0 \implies \beta = -\epsilon\lambda + \frac{3\lambda^2}{16\pi^2}$$

Example:
$$0 \neq 3$$
 in $D=6-\epsilon$ dim.

(39)
$$[\phi^{3}] = 3(\frac{D-2}{2}) \Rightarrow [g] = D - [\phi^{3}] = \frac{6-D}{2} \Rightarrow g_{0} = \mu^{6/2} \frac{[g+dg]}{(1+d^{2})^{3/2}}$$

(40) Where we calculated $g = g \cdot \frac{g^{2}}{(4\pi)^{3}} \left(\frac{1}{e}\right)$ from $g = \mu^{6/2} \left(\frac{g+dg}{(4\pi)^{3}}\right)$
(10) Where we calculated $g = g \cdot \frac{g^{2}}{(4\pi)^{3}} \left(\frac{1}{e}\right)$ from $g = \mu^{6/2} \left(\frac{g+dg}{(4\pi)^{3}}\right)$ (11) $g = \mu^{6/2} \left(\frac{g+dg}{(4\pi)^{3}}\right)$ (12) $g = \mu^{6/2} \left(\frac{g+dg}{(4\pi)^{3}}\right)$ (13) $g = \mu^{6/2} \left(\frac{g+dg}{(4\pi)^{3}}\right)$ (14) $g = \mu^{6/2} \left(\frac{g+dg}{(4\pi)^{3}}\right)$ (15) $g = \mu^{6/2} \left(\frac{g+dg}{(4\pi)^{3}}\right)$ (15) $g = \mu^{6/2} \left(\frac{g+dg}{(4\pi)^{3}}\right)$ (16) $g = \mu^{6/2} \left(\frac{g+dg}{(4\pi)^{3}}\right)$ (17) $g = \mu^{6/2} \left(\frac{g+dg}{(4\pi)^{3}}\right)$ (18) $g = \mu^{6/2} \left(\frac{g+dg}{(4\pi)^{3}}\right)$ (19) $g = \mu^{6/2} \left(\frac{g+dg}{(4\pi)^{3}}\right)$

(41)
$$\mu \frac{d^2g}{dR_{\mu}} = 0 = \frac{\epsilon}{2}g\left(1 - \frac{3}{4}\frac{3^2}{(4\pi)^3}\frac{1}{\epsilon}\right) + \frac{3}{3}\left(1 - \frac{3}{4}\frac{3^2}{(4\pi)^3}\frac{1}{4}\frac{1}{\epsilon}\right)$$

to lowest order
$$\beta = -\frac{\epsilon}{2}g$$

Notice that \$43 <0 in D=6 (the theory is exymptotically free in the (W) we explain this later

Multiple couplings in various dimensions to ell orders.

With multiple couplings, even with non-renormalisable (i.e. irrelevent) as relevant interactions, we can repeat the strategy by definining dinensionless renormalized couplings in tens of which the dimensionful bone coupling are expensed

(43)
$$\int_{0}^{a} \int_{0}^{b} \int_{0}^{b}$$

where $E = \overline{D} - D$, \overline{D} being the integer dim via aim at $(\overline{D} = 4 \text{ in } \phi^4)$ It's thus convenient to pull out the $\mu^{\overline{D}-\Delta o_{m}\overline{D}}$ -factor

(44)
$$g_{n,bere} = \mu \frac{\log \operatorname{iol} \operatorname{dim} f_{g_n} \operatorname{et} b = \overline{b}}{\sum_{i=1}^{n} \Delta g_n(\overline{b})} \cdot \mu \left(g_n(\mu) + \frac{\alpha_1(g)}{\epsilon} + \frac{\alpha_2(g)}{\epsilon^2} + \cdots \right)$$

Since DQ(D) is linear in D.

$$\left(\alpha_{n} \cdot \mathcal{E} + \overline{\mathcal{D}} - \Delta_{\mathcal{O}_{n}}(\overline{\mathcal{D}}) = \alpha_{n}(\overline{\mathcal{D}} - \mathcal{D}) + \overline{\mathcal{D}} - \Delta_{\mathcal{O}_{n}}(\overline{\mathcal{D}}) = \mathcal{D} - \Delta_{\mathcal{O}_{n}}(\mathcal{D})\right)$$

$$\left(\exp_{\mathbf{x}}(\mathbf{x}) = \mathbf{x} + \lim_{n \to \infty} \frac{\partial^{4}}{\partial \mathbf{x}^{2}} + \lim_{n \to \infty} \frac{\partial^{4}}{\partial \mathbf{x}^{2}}$$

(example: x = +1 in $\frac{1}{2}$ in $\frac{1}{2}$ in $\frac{1}{2}$

For maying coupling $\bar{D} - \Delta \sigma_n(\bar{D}) = 0$ where it is positive and negetive for relevant and irrelevant couplings respectively.

L7/15

(45)
$$0 = \mu \qquad \qquad \int_{0}^{\infty} \int_{0}^{\infty}$$

$$\left(\frac{1}{46}\right) \left\{ \begin{array}{l}
P_{n} = -\left[\int_{0}^{m} - \frac{\alpha_{1}^{(n,m)}}{\epsilon} + ...\right] \left[g_{m} + \frac{\alpha_{1}^{(m)}}{\epsilon} + ...\right] \left(\overline{D} - \Delta_{0}(\overline{D}) + \alpha_{m} \epsilon\right) \\
= -\alpha_{n} g_{n} \epsilon - \left(\overline{D} - \Delta_{0}(\overline{D}) \right) g_{n} + \alpha_{m} g_{m} \alpha_{1}^{(n,m)} - \alpha_{n} \alpha_{1}^{(n)} + \frac{1}{\epsilon} \left(...\right) + \left(\frac{1}{2} - \frac{1}{2}\right) + \frac{1}{2} \left(...\right) + \left(\frac{1}{2} - \frac{1}{2}\right) + \frac{1}{2} \left(...\right) + \frac{1}{2}$$

Now, a finite change in μ require a dange in $g_n(\mu)$ to compensate it: must be finite $\Rightarrow \downarrow (\cdots) + \downarrow (\cdots) + \cdots = 0$ all must vanish identically $\Rightarrow \beta$ is finite as $\leftarrow \infty$

$$\beta_{m} = -(5 - \Delta \theta_{n}(5))g_{n} + \alpha_{m}g_{m} \frac{\partial}{\partial g_{m}} - \alpha_{n}a_{1}^{(n)} \qquad (sum over m)$$
All order result!! chessical:

just from tree-level $g_{n,bare}$ in g_{n} g_{n}

Remark: * This is a very universal formula: the B-function in a generic theory in a generic dimension $D=\overline{D}$ * it depends only on the simple poles of in gn, bare higher poles $\frac{a_{m+2}}{\epsilon^{m+2}}$ do not mater!!

check-1

$$\frac{\lambda \phi^4}{4!}$$
 in D-dimensions has $\alpha = 1$ and $\Delta - \Delta_{\phi^4} = 4-D$

$$\beta \phi^{4} = -(4-8)\lambda + \lambda \frac{2}{3}\alpha_{1} - \alpha_{1} = -(4-8)\lambda + \lambda^{2}\frac{2}{3}(e_{1}/\lambda)$$

$$\lambda^{2}\frac{2}{3}(e_{1}/\lambda)$$

$$\beta_{\phi 4}^{1-loop} = -(k-b)\lambda + \frac{3\lambda^2}{16\pi^2}$$

in ogneement w. 1 Eq. (38)

check-2

$$\frac{1}{2}$$
 $\frac{1}{2}$ in D-dim.

$$\Delta = 1/2$$

$$D - \Delta_{\phi 3} = D - \frac{3(5-2)}{2} = \epsilon$$

$$\sqrt{\beta^2} = -\frac{5}{6} + \frac{1}{7} \frac{3}{3} \frac{3}{3} \left(\frac{3}{3} \right)$$

(51)

$$\Rightarrow \begin{cases} 3 + 3 = -\frac{2}{3} + \frac{3}{2} = -\frac{3}{4} + \frac{3}{4} = -\frac{3}{4} =$$

$$\beta_{43} = -\frac{eg}{2} - \frac{3}{4} \cdot \frac{g^3}{(4\pi)^3}$$

$$(an)^3$$

Remark: (a) The fact that V_{ϵ} , E_{ϵ} ,... terms in Eq. (6)

venish, i.e. finiteness of β when $\epsilon - DO$, it's actually powerful because it fixes higher order poles $\alpha_{i=2}^{(n)}$. from $\alpha_{1}^{(n)}$ only.

(b) Take for example a single coupling like in 49-theory
for simplicity and Keep track of o(=1)-terms in RG-equation

(53)
$$0 = \mu^{-\epsilon} \frac{\partial}{\partial k_{m}} \lambda_{\text{sore}} = \epsilon \cdot \lambda \left(1 + \frac{2 \alpha_{i} / \lambda}{\epsilon^{i}} \right) + \beta \left(1 + \frac{2 \alpha_{i} / \lambda}{\epsilon^{i}} \right) + \beta \left(1 + \frac{2 \alpha_{i} / \lambda}{\epsilon^{i}} \right)$$

$$\beta = -\epsilon \cdot \lambda \left(1 + \frac{2 \alpha_{i} / \lambda}{\epsilon^{i}} \right) = -\epsilon \lambda + \lambda^{2} \frac{\partial}{\partial k_{m}} \left(\frac{\partial}{\partial k_{m}} \right) + \frac{\partial}{\partial k_{m}} \left(\frac{\partial}{\partial k_{m}}$$

Drequiring & (...) = 0 set a differential equation for ash

(54)
$$\frac{\partial (e_1/\lambda)}{\partial \lambda} = \frac{\partial (e_1/\lambda)}{\partial \lambda} \frac{\partial a_1}{\partial \lambda} \implies \text{solving for } a_2/\lambda$$

Likewise one determines a_{n} z_{1} too from setting to zero $\frac{1}{\epsilon^{n}}$ z_{1} .

From an -alone one can extract ell am's, it's just a matter of knowing a_{1} $= \left(\frac{3\lambda}{16\pi r}\right) + \frac{1}{2}\left(\frac{\lambda}{16\pi r^{2}}\right)^{2} + \cdots$ to higher loop

For example, in $\frac{4}{4!}$ -theory in D=4, $a_1 = \frac{3\lambda^2}{16\pi^2}$ so that

(55)
$$a_{2}/\lambda = A \left(\frac{\lambda}{16\pi^{2}}\right)^{2} = \frac{69.54}{16\pi^{2}} + 2A \cdot \frac{\lambda}{(16\pi^{2})^{2}} = \frac{3}{(6\pi^{2})} \cdot \frac{3}{(6\pi^{2})} \cdot \frac{3}{(6\pi^{2})} \cdot \frac{3}{(6\pi^{2})} \cdot \frac{3}{(6\pi^{2})} \cdot \frac{3}{(6\pi^{2})} = \frac{3}{(6\pi^{2})} \cdot \frac$$

i.e. we determined az = 9/3/642)2 from az et 1-loop.

This corresponds disynametrially to XX being the square of XX giving rise to Vez (or lan in Pauli-Villand or hart cutoff) whereas χ and \longrightarrow are genuinely new, $O(\frac{\lambda^3}{16\pi^2 x^2}) \stackrel{!}{\in}$, that contribute to a_1 et a_2 The RG guarantees this in generalty. We can also see it at the level of the amplitude:

(c) From Weinbey theorem, the to pole correspond to the leading powers of Log = > the leading Log's are determined!

Example: $H_{2\rightarrow2} = -\lambda \left(1 + \frac{3\lambda}{16\pi^2} \frac{\log \xi_n}{\log \xi_n} + \left(\frac{\lambda}{16\pi^2}\right)^2 \left[\frac{\lambda}{16\pi^2} \frac{\log \xi_n}{2 - \log \xi_n} + \frac{1}{16\pi^2}\right]$ et 2-loop $\frac{\xi_{2}}{16\pi^2} = -\lambda \left(1 + \frac{3\lambda}{16\pi^2} \frac{\log \xi_n}{\log \xi_n} + \frac{1}{16\pi^2}\right)^2 \left[\frac{\lambda}{16\pi^2} \frac{\log \xi_n}{\log \xi_n} + \frac{1}{16\pi^2}\right]$ et 2-loop (56)

 $\Rightarrow \frac{d}{dh} = 0 = -\beta \left(\frac{1}{16\pi^2} + \frac{3\lambda^2}{16\pi^2} + \frac{3\lambda^2}{16\pi^2} \right) + 3\left(\frac{1}{16\pi^2} + \frac{3\lambda^2}{16\pi^2} - \frac{\lambda^3}{16\pi^2} \right) + 3\left(\frac{1}{16\pi^2} + \frac{\lambda^3}{16\pi^2} + \frac{\lambda^3}{16\pi$ (58)

So we see that the running coupling allows to systematically detaining the higher order tog's. This is a general fearture of RG-equations and it's quite important because it allows to keep perturbation theory under control even with exp. separated energy scales

leading log2 dream determined et 1-loop B-fuc.

determined by 2-loop.
R-funct.

Ry-flow and its Asymptotics

Now, that we see that higher powers of large-lag(E/n) can be determined via RG, we can eddness the asymptotic RG-evolution both O

$$E \gg \mu (> m)$$
 or $(m < c) E < c \mu$ $t = -s - t$

In particular, let's consider some observable clas:

(a con be the
$$\mu = E$$
 in the "physical doice of rex. completion)

(53)

Obs = $g^{*}(1 + g^{2} h(E/h) + (g^{2} h(E/h)^{2} + ...)$

Three $g^{*}(1 + g^{2} h(E/h) + (g^{2} h(E/h)^{2} + ...)$

12 <1 would not seem enough for P.T. when log E/n > 1672?

Can we recover control when $(\frac{g^2}{16\pi^2} \ll 1)$ but $(\frac{g^2}{16\pi^2} & \frac{g}{16\pi^2} \approx 1)$? Yes.

The idea is that for any given E we can change the splitting $L_e + L_{c.\tau}$ so that (A|A, A) is dozen with $A \cong E \implies$ no large leganth opposes

Then, by moving by infinitesimal steps $\mu \rightarrow \mu + \delta_{\mu}$ one can dange to mother $\mu \simeq \epsilon$ from the initial one.

In formules:

(60)
$$\frac{dg}{dh_{n}} = \beta (g) \implies h = \int_{A}^{B} \frac{dg}{\beta (g)} + no large logs in it$$

$$\beta = g(\# \frac{g^{2}}{16\pi^{2}} + \# \frac{g^{2}}{16\pi^{2}} + \# \frac{g^{2}}{16\pi^{2}} + \dots)$$

just powers of 8/672

$$\frac{\lambda}{4!}\phi^{\frac{1}{4}}$$
 in $D=4$ —D $P_{\phi^{\frac{1}{4}}}=\frac{3\lambda^{2}}{\frac{1}{16\pi + 2}}$

$$\ln E/\mu = \int_{\mu}^{E} \frac{d\lambda}{3\lambda^{2}/(6\pi^{2})} = \frac{(16\pi^{2})}{3} \cdot \left(\frac{1}{\lambda/\mu} - \frac{1}{\lambda(E)}\right)$$

$$\lambda(E) = \frac{\lambda(n)}{1 - \frac{3}{16\pi^2} \lambda(n) \log(E/n)}$$

As long as $\lambda(E)$ remains smell we can write $H_2 = (E) + O(\lambda(E)) + O(\lambda(E))$

(63)
$$\frac{M}{2 \rightarrow 2} (E \Rightarrow \mu) \simeq -\lambda(E) = \frac{-\lambda}{1 - \frac{3\lambda}{16\pi^2}} \log(E/\mu)$$
Thus, we plant

and, if the log's one small that we can expand it back (just to Me what the RG-equation (53) is resumming)

(64)
$$M_{2-02}(E) = -k \left[1 + \left(\frac{3k}{16\pi^2}\right) \log(E_k) + \left(\frac{3k}{16\pi^2}\right) \log(E_k) + \dots\right]$$

it nicely moteles the A=9 in Eq. (55-58), but with no effort.

The -> () Log Elm) coefficients ere presided to be 3m. Subleading 1 log n com be medicted once B is colculated to next loop order, and so on.

On the other hand, given a finite coupling I at some scale μ , the coupling become large at a finite energy scale ALANDAU

(65)
$$\Lambda_{LANDAU} \sim \exp\left[\frac{16\pi^2}{3\lambda(\mu)}\right] \cdot \mu$$

 μ being the scale of the initial condition $\lambda = \lambda(\mu)$.

The theory is no longer perturbetive at finite energy:
it can't be extrapolated post En 1 canon. \$\int \footnote in D=4 is EFT

So this means that \$60 theory should be rejected as an ETT with cutoff not larger than A canous. In this sense, \$40 theory is not renormal table, and one should expect higher operators, eg. \$50 to the still unwelly call it "particularized" renormal table because the Lendan-pole xale is exponentially large than \$\mu\$.

Conversely, for a fixed $\lambda(\mu)$ at $\mu_0 >> E$, the coupling $\lambda(E)$ goes to zero as $E(\mu_0 - \nabla \theta)$

$$(66) \qquad \lambda(E) \longrightarrow 0$$

$$E_{\mu_0} \longrightarrow 0$$

(in the "continuum limit" à la Welson)

This is called trivillity of \$\phi^4\$ theory if one really tries to send up -D as Keeping \(\lambda\)/\(\rho\)/\(\frac{1}{2}\), as that corresponds from a Wilsonian perspective to try to remove the artiff regulator while Keeping the bare coup. fixed. \(\rightarrow\) on't be defined beyond perturbation thous.

=
$$\frac{\text{Example:}}{3!}$$
 $\frac{3\phi^3}{3!}$ in $D=6$ $\frac{1}{2}$ $\frac{3}{3!}$ $\frac{3}{5!}$ $\frac{3}{5!}$

(65)
$$ln(\xi/n) = \int_{1}^{\epsilon} \frac{d\xi}{d\xi} = \frac{(4\pi)^{3}}{3/4} \cdot \frac{1}{2} \left[\frac{1}{2^{2}(\epsilon)} - \frac{1}{2^{2}/n} \right]$$

$$g^{2}(E) = g^{2}(n)$$

$$1 + \frac{3}{2} \frac{(g^{2}n)}{(4\pi)^{3}} \log (E/n)$$

Because $g^2(E \rightarrow \infty) \rightarrow 0$ there is no obstruction in extrapolating this theory to arbitrary short-distance (i.e. remains the cutoff, or taking the cont. limit) all & Yay-Mills are also esymptotically free theories!

Ediscloimer: ϕ^3 in D=6, contrary to act & x.n. in D=4, does not in fact makes sense fully non-particularized become af vacuum instability... but for a small anough to tunneling rate to a new vacuum is exp. small... We are using of just as to model \Box

Conversely, the coupling become logs in the IR: $g(E - DA_{IR}) \gg (\pi)^3 - o$ form band state?

(C-1) A separated by A!

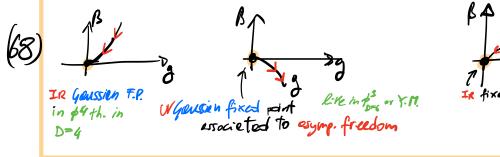
A new IR mess scale linkal to pair (g(A1,A) given in the UV: Dim. transmutation

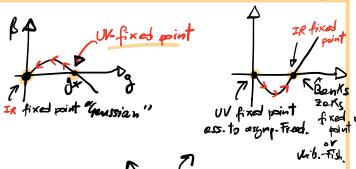
- Fixed Points

es example of asymptotis

We have seen that B>0 leads to coupling growing (develoing) in the UV (IR) and vierese for B<0

(anows point toward IR)





There could be however a situation where 15-00 for some 3-24 If the running mess venish et g=gx, it's collect a Fixed point

(69)
$$\beta(g_{*}) = 0 \qquad \text{when this hoppens } g(u) = g_{*} = const$$

$$(m = 0)$$

(With several couplings
$$\beta_i(g_1,g_2,...)=0$$
 all simultaneously.)

If B has simple zero around gx, \leftarrow lineon in $(g-g_*)$ so g_* can be consider, comparing W1. (47) B=-D-4 $\beta = \beta_* (g - g_*) + \cdots$ around F.P $g = g_*$ or with more couplings $B_i = \frac{\partial B_i}{\partial g} | (g - g +)_i + \cdots$

linear in g-g*, so that comparing to Eq. (7) B=-D-DJg+... we can , (redefining g-gx = g Bg = Bx g), call Bx the quantum or enomelous dimension of a. (lith multiple couplings, in general, need to dieyon of your This is further justified by integrating the RG-flow near gx: [17/1924

$$(71) \quad \ln E/\mu = \int_{g/\mu_1}^{g(e)} \frac{dg}{\beta lg} \frac{1}{g} \int_{g/\mu_1}^{g/\mu_2} \frac{dg}{\beta lg} = \frac{1}{\beta_{1}} \ln \left| \frac{g(e) - g_{1}}{g(\mu) - g_{2}} \right|$$

that is

that is

(72)
$$(g(E) - g_*) = const \cdot (E/\mu)^{B_*}$$

Simple power-law scaling controlled by B_* .

This should be contracted with the free them $B = -(D-\Delta)g$ +... Where the $g_* = 0$ and it's reached by the classical direction set by $D-\Delta$.

If green <= 1 we can epproach the F.P. within perturbation theory.

Exemple: Wilson-Fisher Fixed point in
$$D = 4-E$$
 dimensions

The toy-model thoon we consider now is ϕ^4 theory with a moss term in $D=4-\epsilon$ dimensions with $\epsilon \neq 0$! (in fact in stat. meds. this is expressed to go to $\epsilon \to 1$ to express the 2D Ising model...)

We have two "couplings" (1 m²) eleady calculated $\beta_{\lambda} = -\epsilon \lambda + \frac{31^2}{16\pi^2}$

(73)
$$m_0^2 = \mu^2 g_{2,berc} = \mu^2 (g_2 + \delta g_2) - \mu^2 (g_2 + \delta g_2) - v$$
 need the mass c.T. δg_2

(74)
$$\sum_{\text{self-en.}} = -\frac{\lambda}{4!} \frac{\mu^2 D}{4 \cdot 3} \cdot \int_{\frac{\pi}{(2\pi)^2}} \frac{1}{(\kappa^2 + m^2)} = \frac{\lambda}{16\pi^2} \frac{m^2 L}{E} + \cdots$$

$$(e.g. bs. and first, now the as$$

, a first, ne the eg-div, extract 1, int. beck)

L7/p25

$$\left(\widetilde{W}^{(2)} = \left(\rho^2 + m^2 - \Sigma \right) \Rightarrow$$

$$\left(\widetilde{W}^{(2)} = \left(p^2 + m^2 - \Sigma\right) \implies m^2 = \frac{1}{e} \frac{\lambda}{16\pi^2} m^2 = \left(\frac{1}{e} \frac{\lambda}{16\pi^2} g_2 / \mu^2\right)$$

$$(75) \quad \text{So} \quad \begin{cases} \alpha_1^{(32)} = \frac{1}{16\pi^2} 32 \\ \alpha_2^{(3)} = \frac{1}{16\pi^2} 32 \end{cases}$$

(75) so
$$\begin{cases} a_{1}^{(32)} = \frac{1}{16\pi^{2}} g_{2} \\ a_{1}^{(A)} = \frac{3\lambda}{16\pi^{2}} \lambda \end{cases}$$

$$g_{n} = \begin{pmatrix} g_{2} \\ g_{4} = \lambda \end{pmatrix}, g_{n}, bone = \begin{pmatrix} \mu^{2}(g_{2} + a_{1}^{(2)} + ...) \\ \mu^{\epsilon}(\lambda + a_{n}^{(A)} + ...) \end{pmatrix}$$

and using Eq. (17) (w./ $\alpha_{g_2}=0$, the mass has always the com. dim sine [n]=[0]

$$\beta_{m} = -(\overline{D} - \Delta \theta_{n}(\overline{D}))g_{n} + \alpha_{m}g_{m} \frac{\partial}{\partial g_{m}} - \alpha_{n}\alpha_{2}^{(n)}$$

$$\beta_{m} = -(\overline{D} - \Delta \theta_{n}(\overline{D}))g_{n} + \alpha_{m}g_{m} \frac{\partial}{\partial g_{m}} - \alpha_{n}\alpha_{1}^{(n)}$$

$$\beta_{i} = \begin{pmatrix} \beta_{\lambda} \\ \beta_{g_{2}} \end{pmatrix} = \begin{pmatrix} \epsilon \lambda + \frac{3\lambda^{2}}{16\pi^{2}} \\ -2 g_{2} + \frac{\lambda g_{2}}{16\pi^{2}} \end{pmatrix} = \begin{pmatrix} -\epsilon + \frac{3\lambda}{16\pi^{2}} \\ \frac{g_{2}}{16\pi^{2}} \\ -2 \end{pmatrix} \begin{pmatrix} \lambda \\ \beta_{2} \end{pmatrix}$$

(73)
$$\begin{cases} \beta_{\lambda} = 0 \\ \beta_{2} = 0 \end{cases} = \begin{cases} \lambda_{*} = \frac{16\pi^{2}}{3} \in (-1)^{2} & \text{if } \epsilon = 3/(\pi^{2}) \\ \beta_{2} = 0 \end{cases} = \begin{cases} \lambda_{*} = \frac{16\pi^{2}}{3} \in (-1)^{2} \\ \beta_{2} = 0 \end{cases} = \begin{cases} \lambda_{*} = \frac{16\pi^{2}}{3} \in (-1)^{2} \\ \beta_{2} = 0 \end{cases} = \begin{cases} \lambda_{*} = \frac{16\pi^{2}}{3} \in (-1)^{2} \\ \beta_{2} = 0 \end{cases} = \begin{cases} \lambda_{*} = \frac{16\pi^{2}}{3} \in (-1)^{2} \\ \beta_{2} = 0 \end{cases} = \begin{cases} \lambda_{*} = \frac{16\pi^{2}}{3} \in (-1)^{2} \\ \beta_{2} = 0 \end{cases} = \begin{cases} \lambda_{*} = \frac{16\pi^{2}}{3} \in (-1)^{2} \\ \beta_{2} = 0 \end{cases} = \begin{cases} \lambda_{*} = \frac{16\pi^{2}}{3} \in (-1)^{2} \\ \beta_{2} = 0 \end{cases} = \begin{cases} \lambda_{*} = \frac{16\pi^{2}}{3} \in (-1)^{2} \\ \beta_{2} = 0 \end{cases} = \begin{cases} \lambda_{*} = \frac{16\pi^{2}}{3} \in (-1)^{2} \\ \lambda_{*} = 0 \end{cases} = \begin{cases} \lambda_{*} = \frac{16\pi^{2}}{3} \in (-1)^{2} \\ \lambda_{*} = 0 \end{cases} = \begin{cases} \lambda_{*} = \frac{16\pi^{2}}{3} \in (-1)^{2} \\ \lambda_{*} = 0 \end{cases} = \begin{cases} \lambda_{*} = \frac{16\pi^{2}}{3} \in (-1)^{2} \\ \lambda_{*} = 0 \end{cases} = \begin{cases} \lambda_{*} = \frac{16\pi^{2}}{3} \in (-1)^{2} \\ \lambda_{*} = 0 \end{cases} = \begin{cases} \lambda_{*} = 0 \\ \lambda_{*} = 0 \end{cases} = \begin{cases} \lambda_{*} = \frac{16\pi^{2}}{3} \in (-1)^{2} \\ \lambda_{*} = 0 \end{cases} = \begin{cases} \lambda_{*} = 0 \\ \lambda_{*$$

$$\frac{\partial \mathcal{B}}{\partial \mathcal{B}} = \begin{pmatrix} -\epsilon + \frac{\epsilon}{16\pi^2} + 0 \\ \frac{G_2^{\pm}}{16\pi^2} \end{pmatrix} = \begin{pmatrix} \epsilon & 0 \\ 0 & -2 + \frac{\epsilon}{3} \end{pmatrix}$$

$$= \begin{pmatrix} -\epsilon + \frac{\epsilon}{16\pi^2} + 0 \\ 0 & -2 + \frac{\epsilon}{3} \end{pmatrix}$$

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$$= \begin{pmatrix} -\epsilon + \frac{\epsilon}{3} + 0 \\ 0 & -2 + \frac{\epsilon}{3} \end{pmatrix}$$

$$= \begin{pmatrix} -\epsilon + \frac$$

(79)
$$\Rightarrow$$
 $\beta_{\lambda} \simeq \epsilon (\lambda - \lambda_{*})$

hear

 $WF-F.P.$
 $\beta_{2} = (-2 + \epsilon_{3}) \beta_{2}$

notice that ϕ^4 is marginally relaxant at the

Genssian F.P. in D=4-E (BX=-EX+... 6>0) but it

192 anous toward IR sine mass remain relevant, need to tune it to zero to reach WF-F.P.

The transfer that $\epsilon > 0$

becomes irrelevent et the Wilson-Fisher F.P. (the mess instead remains relevant at both)

The scaling of g_2 in the Gaussian F.P. would be $\beta g_2 = -2g_2$ (that is $g_2(\mu) = g_2(\mu o) (\mu o/\mu)^2$, which is a conficated way of saying $m_{bore}^2 = g_2(\mu) \mu^2 = g_2(\mu_0) \mu_0^2$

At the Wilson-Fisher F.P. instead, using (73), we get

(80)

 $g_{2}(\mu) = g_{2}(\mu_{0})$ μ_{0} μ_{0} μ_{0} the mass is still relevant but less so for $\epsilon > 0$ μ_{0} μ_{0}

More generally, the mess anomolous dim, is defined by taking aut the classical pince

 $-\frac{1}{2} \frac{1}{m^2} \frac{d m^2}{d h_{\mu}} = -\frac{1}{2} \frac{1}{m^2} \frac{d n^2 g_2}{d h_{\mu}} = -\frac{1}{2} \left(2 + \mu^2 \beta_2 \right) = -\frac{1}{2} (\cdots) = -\frac{\epsilon}{6}$ $-2g_2 + (\cdots) = -\frac{\epsilon}{6}$ F.F.

Once one extracted the osymptotic behavior of compling & messes, the esymptotic behavior of other observables is easy to extract. Consider for example, on observebble with a certain most dimension

minimize the effect of large logs by doosing $\mu \simeq E$

(83)
$$Obs = E^{\Delta} F(1, g_i(E), \underline{m(E)}) - D E^{\Delta} [g_i(E) + \sharp g_i(E) + ...]$$

$$no loge log. set leading E-sely if = 1$$

(e.g. woss-section in ϕ^4 th. $\sigma(E\gg E_0) \propto \frac{1}{E^2} \lambda^2(E) = \frac{1}{E^2} \frac{\lambda^2(E_0)}{(1-3\frac{\lambda^2(E_0)h_0^2}{E_0})}$ not just noive VE^2 suching.

Covered: neglecting meso can be on issue when there are IR-div. = IR-div. = IR-sefe observables

Covert: neglecting mess can be on issue when there one IR-div. = Dured to Rook a TR-safe obsaultes

Thicidentally, from the RG equations of couplings & masses one can extract RG-ep. for obsaules

(84)
$$\mu d Obs = 0 = (\mu \frac{2}{2} + \beta_i \frac{2}{2} \frac{1}{2}) Obs = 0$$

if including moses among couplings, otherwise add $-8_m m^2 \frac{1}{2} \frac$

Callen - Symenzik Equation

In addition to dobs =0 or digit =0, dimbere =0 one consider how correlation functions scale with u.

(84)
$$\langle \phi_{0_1} \dots \phi_{0_n} \rangle = \frac{n}{n!} \sqrt{2i} \langle \phi_{1} \dots \phi_{n} \rangle$$
bore fields $\sqrt{\text{enom. correlators}}$

 $\phi_{oi} = \sqrt{2} \phi_{i}$

so that ud bore correlator > = 0 gives

(85)
$$\left(\sum_{i} \delta_{i} + \mu \frac{1}{\delta \mu}\right) \langle \psi_{i} \dots \psi_{n} \rangle = 0$$

$$V_i = \frac{1}{2} \frac{d \ln Z_i}{d \ln \mu}$$

Example: in \$ in D=6, 59. kaf42)

Cellons-Symanzik. equation

(86)
$$\left(\sum_{i} \chi_{i}^{2} + \mu_{2}^{2} + \beta_{i} \sum_{j=1}^{2} (\sum_{j=1}^{2} \chi_{i}^{2}) < \phi_{1} \ldots \phi_{n} > 0$$

egoin including messes anamouply if not sood for fi

X = ½· 6 (€)·9·2·2(-€9+...)

The formel solution of (86) or (85) is obtained by just recolling (84)

$$\langle \phi_1 \dots \phi_n \rangle = \mathcal{G}^{(n)}(\mu, g(\mu), P)$$

(88)
$$G'(\mu, g(\mu), p) = G'(\mu, g(\mu), p) \cdot \frac{T[\sqrt{2i}[g(\mu)]}{T[\sqrt{2i}[g(\mu)])}$$

So we need first to solve for the 2 dependence on μ . Let's work for simplicity with a single field and a single compling:

(89)
$$8 = \frac{1}{2} \frac{d \ln 2}{d \ln n} \implies \ln \frac{2(n)}{2(n_0)} = \int_{n_0}^{\infty} \langle g(n) \rangle d \ln n = \int_{0}^{\infty} \frac{g(n)}{2(n_0)} dg$$

$$(30) \sqrt{2(\mu)} = \sqrt{2(\mu)} \exp\left[\int_{30}^{8} \frac{\chi(y)}{|\lambda(y)|} dy\right]$$

change vanielles

de = B

(31) $G^{(n)}(\mu, g(\mu), \rho) = \exp \left[-n \int_{g_0}^{g_0} \chi(g_0) dg \right] \cdot G^{(n)}(\mu_0, g(\mu), \rho)$ Notice that et a fixed point g = g+ = const so that from (3) 8 = 8 + 3F.P. of g = g + 3 = 5 = 5 = 5 = 5 = 5 = 5 = 5 = 5 = 5 = 6 =(32) (33) $G(n) = (n/p) \cdot G(n, g^*, p)$ $g = g^* \in P$ Scaling correlation at a Fixed Point Example: 2pt - (\$\phi_1 \phi_2 > near a fixed point) From eq. (86) => (u2+28x) 9(x1, x2, µ)=0 Thereise pull out 1/1/2/20 | X-Y2| (|X1-Y2|1/4) + trouslations + dim. and as 1) $(\mu \frac{1}{2} + 2 + \chi +) = \frac{1}{2} (|\chi_1 - \chi_2|_{\mu}) = 0 \Rightarrow \chi = \frac{1}{2} (|\chi_1 - \chi_2|_{\mu}) = 0 \Rightarrow \chi = -2 + \chi = -2$ $G(x) = const \times \frac{-2\delta_{*}}{}$ (35) $G(x_1, x_2, \mu)_1 = \frac{1}{|x_1 - x_2|^2} \cdot \frac{\text{const.}}{|x_1 - x_2| \mu}^2 \times \epsilon$ it scoles like if []=1+4 = const $|x_1 - x_2|^{2(1+\delta_4)}$ justifying the name of V.

One can reach the same conclusion in mon. space from the solution in Eq. (33)

(96)
$$\hat{q}^{(2)} = \hat{q}^{(2)}(\mu, 3*, 9) = (\mu/\mu_0) \hat{q}^{(2)}(\mu_0, g*, 9)$$

in mom. space $\mathcal{P} = (\mu/\mu_0) \hat{q}^{(2)}(\mu_0, g*, 9)$
 $= (\mu/\mu_0) \hat{q}^{(2)} = \hat{q}^{(2)}(\mu_0, g*, 9)$
 $= (\mu/\mu_0) \hat{q}^{(2)} = \hat{q}^{(2)}(\mu_0, g*, 9)$

(37) => F honogeneous finc. degree
$$f_{*}$$
 $F(x) = x^{*}$

(38)
$$\hat{q}^{(2)} = \frac{1}{p^2} \cdot \frac{1}{(p^2 \mu^2)^{-1/4}} = \frac{D}{F.T.} \qquad \frac{1}{(x_1 - x_2)^2 (1 + x_4)}$$

Finely, a slightly more pedestrien derivation based on partials the would be to minimize the large logs.

(33)
$$\frac{70}{4}(\mu, 34, p) = (\mu/\mu_0) \frac{9}{9}(\mu_0, 84, p)$$

$$= (\mu/\mu_0) \frac{1}{9} \left[1 + O(\frac{3}{1671} + \frac{1}{12} + \frac$$